# The Eagle-1 time synchronization scheme

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- Main challenge in satellite-based QKD: frequency drift due to Doppler shift + high loss
- Typical values for LEO satellites:
  - Doppler shift 2e-5, chirp 4e-7 / s
  - 40-60 dB loss









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- Less interfaces = better
  - Don't rely on classical communication
  - Do sync in the quantum channel





- Less interfaces = better
  - Don't rely on classical communication
  - Do sync in the quantum channel
- We need bright pulses for the phase lock anyway [1]
  - $\rightarrow$  use them for the clock recovery in addition!





[1] Bastian Hacker et al 2023 New J. Phys. 25 113007







- Doppler shift + chirp due to satellite orbitWeak signals
- Single photon detector → no intensity information, only timestamps
- Which sent states do the received ones belong to?













#### • Measurements and output:











Time

























- Why do we need t within symbol?
- Quantum pulses have different temporal distribution than dark counts → time window



- What we want to do:
  - Relative sync: frequencies!
    - Determine exact signal rate
  - Absolute sync: offsets!
    - Frame number? ~200 μs
    - Index of quantum pulse? ~400 ps
    - Where is the "middle" of a quantum pulse? ~10 ps





- How to proceed?
- GHz signal rate
- Timing requirement  $\rightarrow$  < 1 s time to recover clock  $\rightarrow$  work on  $\sim t_{total} = 100$  ms blocks
- With reference pulses: < 1 kCounts/block





- Required period accuracy  $\Delta T$ ?
- Period T
  - $\rightarrow$  offset after sifting block  $\ll T$
  - $\rightarrow$  offset after one period:  $\ll T \cdot T/t_{total}$
- Per period T:  $\Delta T < 1 fs$
- Translated into frequency:

$$\Delta F = \frac{\Delta T}{T} \cdot F < 10 \ Hz$$



- FFT?
  - *t<sub>total</sub>* = 100 ms: ~1e9 entries
  - FFTing a 10<sup>9</sup> array with all 0s except for few k 1s?





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MemoryError

# Relative sync

- FFT?
  - Frequency already known to some degree, no need to check every frequency from 0 to 1/T with Hz resolution!
  - Arrays too long / computation too slow
- $\rightarrow$  Epoch folding!





# What is epoch folding?

- Rosenberg et al., 1975
- Astrophysics
- Recent renaissance within the QKD context
  - BayernQSat internal documentation, 2021
  - [1] Spiess and Steinlechner, 2023
  - [2] Zahidy et al., 2023

# letters to nature

Observations of a transient X-ray source with a period of 104 s

Received June 13; accepted July 18, 1975.



An epoch-folding analysis yields the best estimate for the period of 104.14  $\pm$  0.16 s (1 $\sigma$ ). Figure 3 shows the observed



MAX PLANCK INSTITUTE FOR THE SCIENCE OF LIGHT [1]: https://arxiv.org/pdf/2311.14104.pdf [2]: https://arxiv.org/pdf/2212.12589.pdf



# What is epoch folding?

- Expected period *T*<sub>guess</sub>:
  - For each timestamp t, get t % T<sub>guess</sub> and sort into histogram
- Repeat for several *T<sub>guess</sub>*
- Histogram with highest variance belongs to correct period



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Τ?







Τ?

4		•	
3	•		
2	•	•	•
1			•
0	•	•	•







Τ?

4			
3			
2	•		
1			
0			
			1







T  $\bigcirc$ 4 3 2 1 L 0 I 





















- Guess temporal signal distribution
- $\rightarrow$  in period T, how long is the signal phase a\*T?









Bandwith of variance?











- Period array resolution at least  $a \cdot \frac{T^2}{t_{total}}$
- $\rightarrow$  this is also the highest accuracy one can achieve with limited  $t_{total}$  ( $\rightarrow$  "FWHM")
- Exact ratio between  $var(T_{correct})$  and  $var(T_{wrong})$  depends on dark and background counts (=noise)  $n_{dc}$ , amount of clicks  $n_S$  and the temporal shape of the signal.





 Start with low frequencies (frame rate), work forward to high frequencies (signal rate)



 Integer factors between them! → allows to use all recorded data without adding artifacts







# Finding frequencies

- Start with low frequencies (frame rate), work forward to high frequencies (signal rate)
- Higher frequencies are integral multiples of the lower ones







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- Finding frequencies
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- How does epoch folding perform with low count rates and high noise → QKD?
- → Monte Carlo simulations
- As expected: increasing signal to noise ratio or measurement duration decreases failure
   probability



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 Easy, just increase t<sub>total</sub> until failure probility=0!



- Chirp  $\gamma_1$  (LEO orbit!) limits  $t_{total}$   $t_{total} < \sqrt{\frac{2aT}{\gamma_1}} = 20 ms$
- Luckily, strongest chirp during zenith → better
  SNR







# Likelihood of failure





• Chirp  $\gamma_1$  (LEO orbit!) limits  $t_{total}$   $t_{total} < \sqrt{\frac{2aT}{\gamma_1}} = 20 ms$ 



• Luckily, strongest chirp during zenith  $\rightarrow$  better **SNR** 



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# Removing chirp

- Signal rate is time dependent
- Easy approach:
  - divide recorded trace in subtraces
  - Get signal rate  $f_S$  per subtrace
  - Cubic fit  $1/f_S \approx a_0 + a_1 \cdot t + a_2 \cdot t^2 + a_3 \cdot t^3$

$$t_S = \sum_{k=1}^3 t_R^k \cdot a_k / k!$$





- Faster approach: get signal rate only once for whole trace
- Then cut trace into subtraces, get histograms with the signal rate calculated before
- Since frequency drifts slowly,  $\Delta f$  is visible in a phase offset for each subtrace









- Now we can find a function for the temporal offset  $\Phi(t)$
- The doppler shift  $\gamma(t)$  is then defined as

• 
$$\gamma(t) = \frac{d\Phi(t)}{dt}$$
  $t_S = \sum_{k=1}^3 t_R^k \cdot a_k / k!$ 

- Now we can remove the chirp and find a single signal rate for the complete recorded trace
- Relative synchronization







- Relative synchronization
- $\rightarrow$  Absolute sync?

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- $\rightarrow$  Frame numbers, encoded in the reference pulses via early/late
  - 5k frames per second
  - 158 reference pulses within frame every 80 ns
  - First 127 pulses: PRBS sequence
  - Last 31 pulses: frame number











 $\rightarrow$  Bit synchronization with satellite without absolute time synchronization





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empty slot

20 ns

#### Find start of frame via PRBS correlation





# Find offsets!

- Find start of frame via PRBS correlation
- $\rightarrow$  ~10 ns accuracy

Histogram and PRBS sequence (zoomed in)





# Find offsets!

Increase accuracy with higher frequencies



# Find offsets!

- Increase accuracy with higher frequencies
- $\rightarrow$  ps accuracy





- Find offset with PRBS correlation
- Decode information using monotonously increasing bits





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- Find offset with PRBS correlation
- Decode information using monotonously increasing bits
- Very robust, works with 0.01 received photons per bit
   Detected photons per bit: 0.001





- Find offset with PRBS correlation
- Decode information using monotonously increasing bits
- Very robust, works with 0.01 received photons per bit
- Absolute sync

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# Summary

- Relative sync ✓
  - Epoch folding

[Scientific publication in preparation]

- Higher frequencies are integral multiples of lower ones
- Absolute sync
  - Correlations and frame numbers



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#### Thank you for your attention!

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